Experimental study of the $\gamma p \to K^0 \Sigma^+$, $\gamma n \to K^0 \Lambda$, and $\gamma n \to K^0 \Sigma^0$ reactions at the Mainz Microtron

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This work measured $d\sigma/d\Omega$ for neutral kaon photoproduction reactions from threshold up to a c.m. energy of 1855 MeV, focusing specifically on the $\gamma p \to K^0 \Sigma^+$, $\gamma n \to K^0 \Lambda$, and $\gamma n \to K^0 \Sigma^0$ reactions. Our results for $\gamma n \to K^0 \Sigma^0$ are the first-ever measurements for that reaction. These data will provide insight into the properties of N^* resonances and, in particular, will lead to an improved knowledge about those states that couple only weakly to the πN channel. Integrated cross sections were extracted by fitting the differential cross sections for each reaction as a series of Legendre polynomials and our results are compared with prior experimental results and theoretical predictions.

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I. INTRODUCTION

Most of our early knowledge of N^* resonances came from experiments involving the πN channel in the initial or final state, e.g., pion nucleon elastic or inelastic scattering [1] or single-pion photoproduction. Lattice QCD and quark models both predict more nucleon resonances in the mass range below 2000 MeV than have been observed experimentally. This is known as the "missing resonances" problem in baryon spectroscopy. For that reason, there has been a concerted effort at electromagnetic facilities, including JLab, Mainz, and Bonn, to measure N^* formation reactions that do not include the πN chan-

nel at all. The data analyzed in this work bear directly on that problem. The photoproduction of a kaon on a nucleon target can provide new information on nucleon resonances. Out of six elementary kaon photoproduction reactions ($\gamma p \to K^0 \Sigma^+$, $\gamma n \to K^0 \Lambda$, $\gamma n \to K^0 \Sigma^0$, $\gamma p \to K^+ \Lambda$, $\gamma p \to K^+ \Sigma^0$, $\gamma n \to K^+ \Sigma^-$), a significant amount of experimental research [2–6] has been done on the charged kaon reactions.

By contrast, there have been very few published studies of K^0 photoproduction. Lawall et~al. [7] measured $\gamma p \to K^0 \Sigma^+$ at ELSA, in Bonn, using the SAPHIR detector. Events were reconstructed using the $K^0 \to \pi^+\pi^-$, $\Sigma^+ \to \pi^0 p$, and $\Sigma^+ \to \pi^+ n$ decays. Castelijns et~al. [8] performed complementary measurements of $\gamma p \to K^0 \Sigma^+$

at ELSA with events reconstructed using the $K^0 \to \pi^0\pi^0$ and $\Sigma^+ \to \pi^0p$ decays. Aguar-Bartolomé et al. [9] measured $\gamma p \to K^0\Sigma^+$ at Mainz using the Crystal Ball and TAPS detectors with events reconstructed using the $K^0 \to \pi^0\pi^0$ and $\Sigma^+ \to \pi^0p$ decays. Recently, Compton et al. [10] measured $\gamma n \to K^0\Lambda$ at JLab using the CLAS detector. Data were collected in two datasets, g10 and g13, which used different run conditions. Events were reconstructed using the $K^0 \to \pi^+\pi^-$ and $\Lambda \to \pi^-p$ decays.

The main focus of the current work was to measure the differential cross section from threshold to c.m. energy W=1855 MeV for the reactions $\gamma p \to K^0 \Sigma^+$, $\gamma n \to K^0 \Lambda$, and $\gamma n \to K^0 \Sigma^0$ on a liquid deuterium target, where W was calculated from the incident beam energy assuming quasifree kinematics. The measurements were performed at MAMI-C, the Mainz Microtron located in Mainz, Germany. We analyzed these reactions via the $K^0 \to \pi^0 \pi^0$ decay. Further details are provided in Sec. III.

The cross-section data can be used to help determine N^* resonance properties using partial-wave analyses or to test phenomenological models of kaon photoproduction. This paper reports the world's first results on differential and total cross sections for the reaction $\gamma n \to K^0 \Sigma^0$.

This paper is divided into six sections: Sec. II describes the experimental setup, Sec. III describes the data analysis, Sec. IV describes the calculation of uncertainties, Sec. V describes the results and discussion for all three reactions, and Sec. VI gives the summary and conclusions. Our measured cross sections are tabulated in the appendix.

II. EXPERIMENTAL SETUP

Data for the photoproduction of neutral kaon reactions on a liquid deuterium target were measured using the Crystal Ball (CB) [11–15], particle identification detector (PID) [16] and TAPS [13-15] detectors. All these detectors were set up at the Mainz Microtron [17] bremsstrahlung-tagged photon beam facility in Germany. At the time the measurements were performed, MAMI-C could deliver electrons with energies up to a maximum energy of 1508 MeV. The mono-energetic electron beam was used to produce photons via bremsstrahlung in a 10- μ m copper radiator. The bremsstrahlung photons are tagged by the Glasgow photon tagger [18]. The tagged photons are then passed through a lead collimator to produce a photon beam. The hole in the lead collimator was 4 mm in diameter for this experiment. This collimation gave a photon beam spot on target with a diameter of about 1.3 cm. The photon beam was incident on a $125 \mu m$ Kapton target cylinder of length 4.72 cm and diameter 4 cm. Further details on the target system can be found in Ref. [19].

The Crystal Ball (CB) is a multiphoton spherical spectrometer [11]. The CB geometry is based on an icosahedron, a polyhedron having 20 triangle-shaped sides. Each

of the 20 major triangles is divided into four minor triangles. Each minor triangle consists of nine crystals, so for a complete sphere, there would be 720 crystals. However, for the entrance and exit tunnels, 48 crystals were not installed, resulting in 672 crystals for the Crystal Ball. The chemical composition of each crystal is thallium-doped sodium iodide, NaI(Tl), which is a hygroscopic material so it is important to protect the crystals from moisture [12, 20]. The Crystal Ball covers the polar angle range from 20° to 160° and the azimuthal angle range from 0° to 360°.

The forward moving particles are detected by TAPS [12, 21], which was configured as a photon calorimeter consisting of $384~\mathrm{BaF_2}$ crystals located downstream of the Crystal Ball. These $\mathrm{BaF_2}$ crystals were arranged in a honeycomb pattern to form a hexagonal wall covering the polar angle range from 4° to 20° .

The PID (Particle Identification Detector) [16] is a cylindrical detector with a 5-cm inner radius oriented concentric with the target inside the Crystal Ball. It was designed to work along with the CB to provide information on charged particles. The PID distinguishes between different types of charged particles and neutral particles based on the energy deposited in the PID elements versus total energy measured in a CB cluster. For further details about these detectors, such as their energy and angle resolutions or their calibrations, see [13–15, 20, 22– 25]. The CB and TAPS detectors are very efficient at detecting the final-state photons. A cylindrical MWPC (MultiWire Proportional Chamber) may be used to improve the angular resolution (tracking) of charged particles. During this experiment, the MWPC was not used. Figure 1 shows a schematic diagram of the CB and TAPS detector setup.

III. DATA ANALYSIS

After all the detectors had been calibrated, the event selection and analysis was carried out. Detailed Monte Carlo (MC) studies were performed using 3×10^6 events generated according to phase space for each of the three K^0 photoproduction reactions, as well as for $\gamma p \to \eta p$ and $\gamma n \to \eta n$, which are the leading backround reactions due to $\eta \to 3\pi^0 \to 6\gamma$ decays.

In each reaction the K^0 was identified through its decay $K^0 \to \pi^0 \pi^0 \to 4\gamma$. The Σ^+ was identified through its decay $\Sigma^+ \to \pi^0 p$, Λ through its decay $\Lambda \to \pi^0 n$, and Σ^0 through its decay $\Sigma^0 \to \gamma \Lambda \to \gamma \pi^0 n$. Therefore, the detection of three π^0 s in the final state was required in all cases, giving rise to six final-state photons via $\pi^0 \to \gamma \gamma$. Data for $\gamma p \to K^0 \Sigma^+$, $\gamma n \to K^0 \Lambda$, and $\gamma n \to K^0 \Sigma^0$ reactions were sorted into various cases (nc), where n represents the detected number of final-state neutral particles and c represents the detected number of final-state charged particles. Table I tabulates the reactions and the corresponding cases for the present work.

TABLE I.	Cases	based	on	nucleon	detection	for	all	three
$\gamma N \to K^0 Y$	<i>r</i> .							

Case	Reaction	Comment
61	$\gamma p \to K^0 \Sigma^+$	final p detected
60	$\gamma p o K^0 \Sigma^+$	final p not detected
	$\gamma n \to K^0 \Lambda$	final n not detected
70	$\gamma n o K^0 \Lambda$	final n detected
	$\gamma n \to K^0 \Sigma^0$	final n not detected
80	$\gamma n \to K^0 \Sigma^0$	final n detected

If only six neutral clusters are detected, the event is case (60). To be a viable event for $\gamma p \to K^0 \Sigma^+$ or $\gamma n \to K^0 \Lambda$, further analysis was needed to establish these six neutral clusters as photons produced from π^0 decays. The data analysis for case (60) starts by first selecting events that have six and only six neutral clusters. If the final proton in $\Sigma^+ \to \pi^0 p$ is detected then there will be six neutral clusters and one charged cluster in the final state, which defines case (61). If the neutron in $\Lambda \to \pi^0 n$ is detected then there will be seven neutral clusters and no charged cluster, which defines case (70).

For $\gamma n \to K^0 \Sigma^0$ events, the detection of seven photon candidates is required, six coming from π^0 decays and one coming from $\Sigma^0 \to \gamma \Lambda$. If the final-state neutron is not detected, then the event corresponds to case (70); however, if the final-state neutron is detected, then the event corresponds to case (80).

Once events had been separated according to the number of neutral and charged clusters, the next step was to identify the final three π^0 s from the neutral clusters. To identify the three π^0 s, all distinct possible combinations

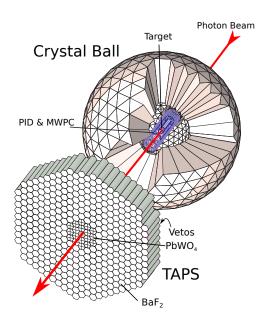


FIG. 1. Schematic diagram of the CB and TAPS detectors. The PID is placed inside the CB for charged particle detection. In this experiment, the $PbWO_4$ crystals were not installed in TAPS.

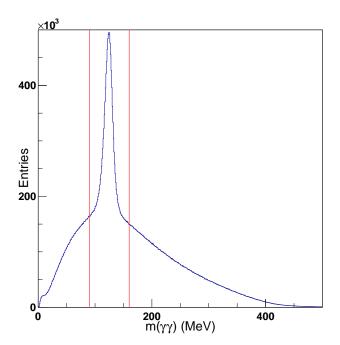
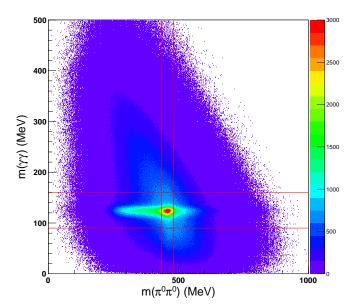
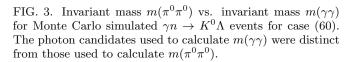


FIG. 2. Invariant mass of all distinct $\gamma\gamma$ combinations for Monte Carlo simulated $\gamma n \to K^0 \Lambda$ events for case (60). The peak corresponds to the π^0 meson. Combinations between the cuts denoted by the vertical red lines correspond to π^0 candidates.

of two-photon candidates were constructed. There are 15, 21, and 28 possible ways to construct distinct two- γ combinations from six, seven, and eight neutral clusters, respectively. A histogram of the invariant-mass of all distinct two- γ combinations for case (60) is shown in Fig. 2. Only those distinct two- γ combinations whose invariantmass $m(\gamma \gamma)$ was between 90 and 160 MeV are the actual π^0 candidates. This invariant-mass cut is represented by solid red vertical lines in Fig. 2. A typical event had several combinations that satisfied this criterion. Only those events that had a minimum of three distinct π^0 candidates were kept. Major sources of background for the reactions of interest are $\gamma p \rightarrow \eta p$ and $\gamma n \rightarrow \eta n$, where $\eta \to 3\pi^0$. In order to eliminate this background, only those three π^0 candidates whose combined invariant mass is greater than 600 MeV were selected for further analysis [9, 26]. This cut significantly reduces the η background contribution while only slightly reducing events from the reactions of interest. If the three π^0 candidates for a given combination are labeled as π_1^0 , π_2^0 , π_3^0 , then there are three ways to construct the two π^0 s that could correspond to a K^0 decay; that is, $(\pi_1^0 \pi_2^0)$, $(\pi_2^0 \pi_3^0)$, or $(\pi_1^0\pi_3^0)$. A histogram of the mass of one π^0 candidate $m(\gamma\gamma)$ versus the invariant mass $m(\pi^0\pi^0)$ of the other two π^0 candidates is shown in Fig. 3. This twodimensional plot provided information on where best to impose a cut on $m(\pi^0\pi^0)$ to reduce the background further. Only combinations in which $m(\pi^0\pi^0)$ was between





435 and 482 MeV were selected for further analysis. This cut was applied before the energy correction discussed below. After this correction, the K^0 peaks in the $\pi^0\pi^0$ invariant-mass distribution were very close to 498 MeV.

Figure 4 shows a histogram of the invariant mass $m(\pi^0 n)$ plotted versus the invariant mass $m(\pi^0 \pi^0)$. The quantity $m(\pi^0 n)$ was actually calculated as the missing mass of the same $\pi^0\pi^0$ combination, since the two quantities should be equal. This plot provided information on where best to impose a cut on the invariant mass $m(\pi^0 n)$. Only combinations in which $m(\pi^0 n)$ was between 1000 and 1300 MeV were selected for further analysis. After the energy correction mentioned above and described below, the peaks in the $m(\pi^0 n)$ distributions were very close to the Λ mass (1116 MeV) for the MC simulated $\gamma n \to K^0 \Lambda$ events. Monte Carlo studies on the polar angle of the undetected nucleon showed that most of the undetected nucleons go forward at our kinematics. A cut was therefore imposed that the cosine of the polar angle of the final-state nucleon, whether measured or calculated, must be greater than or equal to 0.7. All these cuts were used to reduce the number of incorrect three π^0 combinations. Even after all these cuts, there were still a number of events with more than one candidate for the correct three- π^0 combination. Monte Carlo studies were made of the opening angle between two photons for $\pi^0 \to \gamma \gamma$ decays. While the distribution is broad, it is more likely for our kinematics that the opening angle is less than 90° than greater than 90°. The average opening angle for each remaining three π^0 combination was therefore calculated and the combination with the

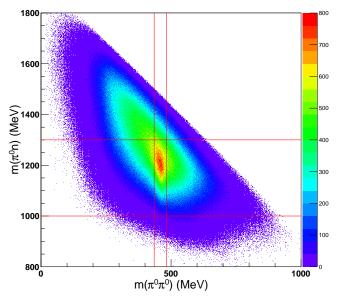


FIG. 4. Invariant mass $m(\pi^0 n)$ vs. invariant mass $m(\pi^0 \pi^0)$ for Monte Carlo simulated $\gamma n \to K^0 \Lambda$ events for case (60). (See text for details.)

minimum average opening angle was selected as the best choice for the correct $3\pi^0$ combination. Although several methods for reconstructing the $3\pi^0$ combination were investigated using Monte Carlo simulations, this method produced the largest K^0 yields.

For case (61), events with six neutral clusters and one charged cluster were selected. The PID was used to select the proton candidate. Similar analysis steps were used to select the best choice for the correct three- π^0 combination as for case (60).

For case (70), there were seven neutral clusters. Similar analysis steps were followed as for case (60) to identify the best choice for the correct three- π^0 combination. Here for each three- π^0 combination there was one unpaired particle.

For case (80), there were eight neutral clusters. Again, similar analysis steps were followed as for case (60) to identify the best choice for the correct three π^0 combination. Here for each three- π^0 combination there were two unpaired particles; i.e, the seventh and eighth particles (a photon and a neutron). The missing mass of the seven photons should equal the mass of the neutron. Therefore a cut was imposed that the missing mass of the three π^0 s and the seventh particle (a photon) be greater than 800 MeV and a cut that cosine of the polar angle of the eighth particle (a neutron) should be greater than or equal to 0.7. These cuts were used to reduce the number of incorrect three- π^0 combinations for case (80), and helped to distinguish which of the other neutral particles was a neutron.

The energy reconstruction of the K^0 mesons was im-

proved by applying a correction [20],

$$E' = E \cdot \frac{m_{\pi^0}}{m_{\gamma\gamma}},\tag{1}$$

which made use of information obtained from the good angular resolution of the CB, after the best choice for the correct three- π^0 combination had been determined. Here E is the relativistic energy of each π^0 , $m_{\gamma\gamma}$ is the invariant mass of the decay photons, and $m_{\pi^0} = 135 \text{ MeV}$ is the known π^0 mass. Before scaling, the invariant mass for $\pi^0 \to \gamma_1 \gamma_2$ is given by

$$(m_{\gamma\gamma})^2 = 2E_1 E_2 (1 - \cos \theta_{\gamma\gamma}), \tag{2}$$

where $\theta_{\gamma\gamma}$ is the measured opening angle for $\pi^0 \to \gamma_1 \gamma_2$. Here E_1 and E_2 are the measured energies of the two photon clusters. After scaling $(E_1 \to E_1' \text{ and } E_2 \to E_2')$, the scaled invariant mass $m(\gamma\gamma)$ was exactly the π^0 mass, 135 MeV. The scaled 4-momenta of the π^0 s were used to calculate $m(\pi^0\pi^0)$ and $m(\pi^0N)$, where N represents the nucleon. All three $\pi^0\pi^0$ combinations were considered for further analysis. In MC simulations for each $\gamma N \to K^0 Y$ event, there are two incorrect $\pi^0\pi^0$ combinations for every correct combination corresponding to $K^0 \to \pi^0\pi^0$. In real data, there can be additional contributions to background in the $m(\pi^0\pi^0)$ distributions.

The $\pi^0\pi^0$ invariant-mass distributions were fitted using a binned likelihood method with the parametrization

$$y(x) = \left[\frac{x^2 - (270)^2}{x^2}\right]^{\alpha} \left[\beta \exp\left(-\frac{1}{2}\left(\frac{x - \mu}{\sigma_B}\right)^2\right) + \delta \exp\left(-\frac{1}{2}\left(\frac{x - 498}{\sigma_K}\right)^2\right)\right],\tag{3}$$

where α , β , δ , μ , σ_B , and σ_K were fitting parameters. The first factor ensured that the distribution goes to zero when $x = 2m_{\pi^0} = 270$ MeV. The exponent α is a small number $(0 < \alpha < 1)$ determined by fitting the $m(\pi^0\pi^0)$ distribution for given energy bins. The parameter β measures the yield of the background contribution. The background was represented by a scaled Gaussian distribution with centroid μ and standard deviation σ_B . The parameter δ measures the yield of the kaon signal. The kaon signal distribution was represented by a scaled Gaussian with centroid 498 MeV (the K^0 mass) and standard deviation σ_K . The observed $m(\pi^0\pi^0)$ distributions for each energy bin, summed over all angle bins, were fitted to determine α and σ_K parameter values for each energy bin. Next the observed $m(\pi^0\pi^0)$ distributions for each angle bin, for a particular energy bin, were fitted with the values of α and σ_K held fixed at their fitted values for that particular energy bin. Monte Carlo simulations were used to verify that this approximation was reasonable. The fitting parameters β , δ , μ , and σ_B were allowed to vary freely in each angle and energy bin. The fitted value of μ for a particular angle and energy bin, with α and σ_K held fixed as described above, was called the nominal background centroid. The values of the nominal background centroid for each energy and angle bin

were recorded for further analysis. The background contribution was obtained after the fit by setting δ equal to zero. Numerical integration was used to calculate the total number of kaons (the kaon yield, N_{K^0}) by subtracting the areas under the total and background curves.

The kaon yield was sensitive to the background contribution. A second fit of the observed $m(\pi^0\pi^0)$ distributions was performed with a different value of μ called the modified centroid. The modified centroid was chosen to be the average of the nominal centroid of the background and the signal centroid (498 MeV). This modified centroid was the maximum value of the background centroid that produced a good fit of the data. The use of these two background centroids is discussed further in Sec. IV. Figure 5 shows the observed $\pi^0\pi^0$ invariant-mass distributions for $\gamma p \to K^0\Sigma^+$, $\gamma n \to K^0\Lambda$, and $\gamma n \to K^0\Sigma^0$ summed over all energy and angle bins. The fitted total invariant-mass distributions are represented by solid red curves and the background contributions are represented by solid black curves.

For calculating the differential cross sections, eight angle bins were used to cover the range from $\cos\theta_{\rm cm} = -1.0$ to +1.0. The c.m. energy range W=1615 to 1765 MeV was divided into five bins of width 30 MeV, and the c.m. energy range W=1765 to 1865 MeV was divided into five bins of width 20 MeV. After subtracting the background, the differential cross section for a specified energy-angle bin was calculated using

$$\frac{d\sigma}{d\Omega} = \frac{N_{K^0}}{N_{\gamma} \epsilon N_{\rm t} B 2\pi \Delta \cos \theta_{\rm cm}},\tag{4}$$

where $N_{K^0} = N_{K^0}(E_{\gamma}, \theta_{\rm cm})$ is the kaon yield for a given energy-angle bin, $N_{\gamma} = N_{\gamma}(E_{\gamma})$ is the photon flux for a given energy bin, $\epsilon = \epsilon(E_{\gamma}, \theta_{\rm cm})$ is the acceptance for a specified energy-angle bin calculated from Monte Carlo simulations, $N_{\rm t}$ is the number of target nucleons per cm², B is a product of branching ratios for the particular reaction, and $\Delta\cos\theta_{\rm cm}$ is the bin width for $\cos\theta_{\rm cm}$.

The differential cross section for $\gamma n \to K^0 \Sigma^0$ for case (80) was calculated using

$$\left(\frac{d\sigma}{d\Omega}\right)_{\gamma n \to K^0 \Sigma^0}^{80} = \frac{N_{K^0}^{80}}{N_{\gamma} \epsilon_{\Sigma^0}^{80} N_{\rm t} B_{\Sigma^0} 2\pi \Delta \cos \theta_{\rm cm}}, \quad (5)$$

where $N_{K^0}^{80}$ is the measured K^0 yield for case (80) and $B_{\Sigma^0}=0.05301\pm0.00074.$

For case (70), the measured K^0 yield has contributions from both $\gamma n \to K^0 \Sigma^0$ and $\gamma n \to K^0 \Lambda$: $N_{K^0}^{70} = N_{\Lambda}^{70} + N_{\Sigma^0}^{70}$. Since $(d\sigma/d\Omega)_{\gamma n \to K^0 \Sigma^0}^{70} = (d\sigma/d\Omega)_{\gamma n \to K^0 \Sigma^0}^{80}$,

$$\begin{split} N_{K^0}^{70} &= \left(\frac{d\sigma}{d\Omega}\right)_{\gamma n \to K^0 \Lambda}^{70} \times N_{\gamma} \epsilon_{\Lambda}^{70} N_{\rm t} B_{\Lambda} 2\pi \Delta \cos \theta_{\rm cm} \\ &+ \left(\frac{d\sigma}{d\Omega}\right)_{\gamma n \to K^0 \Sigma^0}^{80} \times N_{\gamma} \epsilon_{\Sigma^0}^{70} N_{\rm t} B_{\Sigma^0} 2\pi \Delta \cos \theta_{\rm cm}, \end{split} \tag{6}$$

where $B_{\Lambda} = B_{\Sigma^0} = 0.05301 \pm 0.00074$. Values of B_{Λ} , B_{Σ^0} , and B_{Σ^+} were calculated using branching ratios

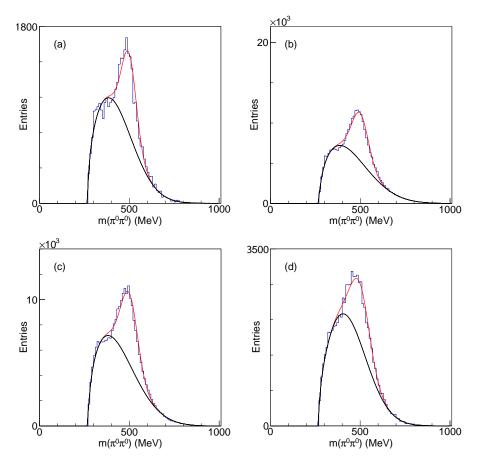


FIG. 5. Observed $\pi^0\pi^0$ invariant-mass distributions: (a) $\gamma p \to K^0\Sigma^+$ for case 61, (b) combination of $\gamma p \to K^0\Sigma^+$ and $\gamma n \to K^0\Lambda$ for case 60, (c) combination of $\gamma n \to K^0\Lambda$ and $\gamma n \to K^0\Sigma^0$ for case 70, (d) $\gamma n \to K^0\Sigma^0$ for case 80. Data in the histograms were summed over all energy and angle bins. The fitted total invariant-mass distributions are represented by solid red curves and the background contributions are represented by solid black curves.

taken from the Review of Particle Physics [27]. For details, see Ref. [28]. The $\gamma n \to K^0 \Lambda$ differential cross section for case (70) is then

$$\left(\frac{d\sigma}{d\Omega}\right)_{\gamma n \to K^{0}\Lambda}^{70} = \frac{N_{K^{0}}^{70}}{N_{\gamma} \epsilon_{\Lambda}^{70} N_{t} B_{\Lambda} 2\pi \Delta \cos \theta_{cm}} - \frac{\epsilon_{\Sigma^{0}}^{70}}{\epsilon_{\Lambda}^{70}} \cdot \frac{B_{\Sigma^{0}}}{B_{\Lambda}} \cdot \left(\frac{d\sigma}{d\Omega}\right)_{\gamma n \to K^{0}\Sigma^{0}}^{80}.$$
(7)

The measured $\gamma n \to K^0 \Sigma^0$ cross sections for case (80) and the measured K^0 yields for case (70) were used to calculate the $\gamma n \to K^0 \Lambda$ cross sections for case (70).

Similarly, the differential cross section for $\gamma p \to K^0 \Sigma^+$ for case (61) was calculated using

$$\left(\frac{d\sigma}{d\Omega}\right)_{\gamma p \to K^0 \Sigma^+}^{61} = \frac{N_{K^0}^{61}}{N_{\gamma} \epsilon_{\Sigma^+}^{61} N_{\rm t} B_{\Sigma^+} 2\pi \Delta \cos \theta_{\rm cm}}, \quad (8)$$

where $N_{K^0}^{61}$ is the measured K^0 yield for case (61) and $B_{\Sigma^+}=0.07637\pm0.00046$.

For case (60), the measured K^0 yield has contributions from both $\gamma n \to K^0 \Lambda$ and $\gamma p \to K^0 \Sigma^+$: $N_{K^0}^{60} = N_{\Sigma^+}^{60} +$

$$\begin{split} N_{\Lambda}^{60} &. \text{ Since } (d\sigma/d\Omega)_{\gamma p \to K^0 \Sigma^+}^{60} = (d\sigma/d\Omega)_{\gamma p \to K^0 \Sigma^+}^{61}, \\ N_{K^0}^{60} &= \left(\frac{d\sigma}{d\Omega}\right)_{\gamma n \to K^0 \Lambda}^{60} \times N_{\gamma} \epsilon_{\Lambda}^{60} N_{\text{t}} B_{\Lambda} 2\pi \Delta \cos \theta_{\text{cm}} \\ &+ \left(\frac{d\sigma}{d\Omega}\right)_{\gamma p \to K^0 \Sigma^+}^{61} \times N_{\gamma} \epsilon_{\Sigma^+}^{60} N_{\text{t}} B_{\Sigma^+} 2\pi \Delta \cos \theta_{\text{cm}}. \end{split} \tag{9}$$

Thus,

$$\left(\frac{d\sigma}{d\Omega}\right)_{\gamma n \to K^{0}\Lambda}^{60} = \frac{N_{K^{0}}^{60}}{N_{\gamma} \epsilon_{\Lambda}^{60} N_{t} B_{\Lambda} 2\pi \Delta \cos \theta_{cm}} - \frac{\epsilon_{\Sigma^{+}}^{60}}{\epsilon_{\Lambda}^{60}} \cdot \frac{B_{\Sigma^{+}}}{B_{\Lambda}} \cdot \left(\frac{d\sigma}{d\Omega}\right)_{\gamma p \to K^{0}\Sigma^{+}}^{61}.$$
(10)

The measured K^0 yields for case (60) and the results of a 15-parameter global fit of $d\sigma/d\Omega$ for $\gamma p \to K^0 \Sigma^+$, discussed in Sec. V.A, were used to calculate the $\gamma n \to K^0 \Lambda$ cross sections for case (60). It was not possible to determine meaningful values of $(d\sigma/d\Omega)_{\gamma p \to K^0 \Sigma^+}^{60}$ due to the large subtractions required.

The final task was to determine the $\gamma n \to K^0 \Sigma^0$ cross section for case (70). For this case, recall that $N_{K^0}^{70} =$

 $N_{\Lambda}^{70} + N_{\Sigma^0}^{70}$. Since $(d\sigma/d\Omega)_{\gamma n \to K^0 \Lambda}^{70} = (d\sigma/d\Omega)_{\gamma n \to K^0 \Lambda}^{60}$,

$$N_{K^{0}}^{70} = \left(\frac{d\sigma}{d\Omega}\right)_{\gamma n \to K^{0}\Lambda}^{60} \times N_{\gamma} \epsilon_{\Lambda}^{70} N_{t} B_{\Lambda} 2\pi \Delta \cos \theta_{cm} + \left(\frac{d\sigma}{d\Omega}\right)_{\gamma n \to K^{0}\Sigma^{0}}^{70} \times N_{\gamma} \epsilon_{\Sigma^{0}}^{70} N_{t} B_{\Sigma^{0}} 2\pi \Delta \cos \theta_{cm}.$$

$$(11)$$

Thus,

$$\left(\frac{d\sigma}{d\Omega}\right)_{\gamma n \to K^0 \Sigma^0}^{70} = \frac{N_{K^0}^{70}}{N_{\gamma} \epsilon_{\Sigma^0}^{70} N_{t} B_{\Sigma^0} 2\pi \Delta \cos \theta_{cm}} - \frac{\epsilon_{\Lambda}^{70}}{\epsilon_{\Sigma^0}^{70}} \cdot \frac{B_{\Lambda}}{B_{\Sigma^0}} \cdot \left(\frac{d\sigma}{d\Omega}\right)_{\gamma n \to K^0 \Lambda}^{60}.$$
(12)

The average of the differential cross sections for the cases with and without detection of the final-state neutron, weighted according to the statistical uncertainties, was calculated for $\gamma n \to K^0 \Lambda$ and $\gamma n \to K^0 \Sigma^0$ and then integrated cross sections were obtained by fitting these values with two-parameter expansions in Legendre polynomials. The Legendre fits include P_0 and P_1 terms for the $\gamma n \to K^0 \Lambda$ and $\gamma n \to K^0 \Sigma^0$ results but just a P_0 term for the $\gamma p \to K^0 \Sigma^+$ results. We used only our case (61) results for $\gamma p \to K^0 \Sigma^+$.

IV. CALCULATION OF UNCERTAINTIES

There are two types of uncertainty involved in calculating the differential cross section. One is the statistical uncertainty and the other is the systematic uncertainty. The statistical uncertainty describes our imprecise knowledge of the kaon signal yield. The systematic uncertainty is the combination of uncertainties from the photon flux, acceptance, and branching ratios. The kaon signal yield in real data was correlated with the centroid of the background. As mentioned earlier, the $\pi^0\pi^0$ invariant-mass distributions were fitted with a sum of scaled Gaussians, with background and signal parts. First the invariantmass histogram was fitted, the background centroid was noted and the kaon yield was calculated; this is called the nominal case. Next a centroid for the background was chosen, which is an average of nominal case background centroid and kaon signal centroid (498 MeV), and the $m(\pi^0\pi^0)$ distribution was refitted and the kaon yield was recalculated. This is called the modified case. The statistical uncertainty was conservatively calculated using

$$\Delta N_{K^0} = [(\text{Poisson error})^2 + (\text{model error})^2]^{\frac{1}{2}}.$$
 (13)

Here, Poisson error = $\sqrt{N_{K^0}+1}$, where N_{K^0} is the average number of K^0 s determined by fitting the $m(\pi^0\pi^0)$ distributions using the nominal and modified values for the background centroid. The model error was taken as the difference in the number of K^0 s determined using the two different background centroids. The statistical

uncertainty in $d\sigma/d\Omega$ is given by

$$\Delta \left(\frac{d\sigma}{d\Omega}\right)_{\rm stat.} = \frac{d\sigma}{d\Omega} \times \left(\frac{\Delta N_{K^0}}{N_{K^0}}\right) \tag{14}$$

and the systematic uncertainty is given by

$$\Delta \left(\frac{d\sigma}{d\Omega}\right)_{\text{sys.}} = \frac{d\sigma}{d\Omega} \times \left[\left(\frac{\Delta N_{\gamma}}{N_{\gamma}}\right)^{2} + \left(\frac{\Delta \epsilon}{\epsilon}\right)^{2} + \left(\frac{\Delta B}{B}\right)^{2} \right]^{\frac{1}{2}}, \tag{15}$$

where the contribution from the uncertainty in the photon flux varied from 1.1% to 2.4% and the contribution from the acceptance varied from about 2% to about 4% for $\gamma n \to K^0 \Lambda$ and $\gamma n \to K^0 \Sigma^0$. The contribution from the product of branching ratios was 1.4% for $\gamma n \to K^0 \Lambda$ and $\gamma n \to K^0 \Sigma^0$ and was 0.6% for $\gamma p \to K^0 \Sigma^+$.

V. RESULTS AND DISCUSSION

A.
$$\gamma p \to K^0 \Sigma^+$$

Figure 6 shows the differential cross section for $\gamma p \rightarrow$ $K^0\Sigma^+$ for the eight energy bins. Our results are shown as solid black circles. Prior results from Lawall et al. [7], measured with the SAPHIR detector at ELSA in Bonn, are shown as solid magenta squares. Prior results from Castelijns et al. [8], measured with the Crystal Barrel and TAPS spectrometers at ELSA, are shown as solid blue triangles. The most precise prior results are from Aguar-Bartolomé et al. [9], measured on a liquid hydrogen target with the Crystal Ball and TAPS spectrometers at MAMI, and shown as solid red circles. Our differential cross-section results are in fair agreement within error bars with prior results in the $\cos \theta_{\rm cm}$ range from +0.6 to -0.45. Our results in the bins at $\cos \theta_{\rm cm} = \pm 0.875$ and -0.675 were unreliable, due to the low statistics and low acceptance at these angle bins. Therefore, those results are not shown in Fig. 6, nor were they used to calculate the integrated cross sections. The solid blue curves in Fig. 6 are from a 15-parameter global fit to all the data. which is described below. The solid red curves are from a three-parameter global fit in which the angular distributions were approximated as being isotropic in each energy bin. The measurements in Fig. 6 are compared with isobar-model predictions by Mart [29], which are shown as dashed green curves. In general, these predictions do not agree well with the measured angular distributions.

In order to ensure a smooth variation with energy and that the cross section vanishes at threshold, a 15parameter global fit of our results and prior differential cross-section data was performed. This fit used the parametrization

$$\frac{d\sigma}{d\Omega} = \sum_{n=1}^{3} \sum_{\ell=0}^{4} a_{n\ell} (W - W_T)^n P_{\ell}(\cos \theta_{\rm cm}), \qquad (16)$$

where $W_T=1687~{\rm MeV}$ is the threshold energy for $\gamma p\to K^0\Sigma^+$ and $P_\ell(\cos\theta_{\rm cm})$ is a Legendre polynomial.

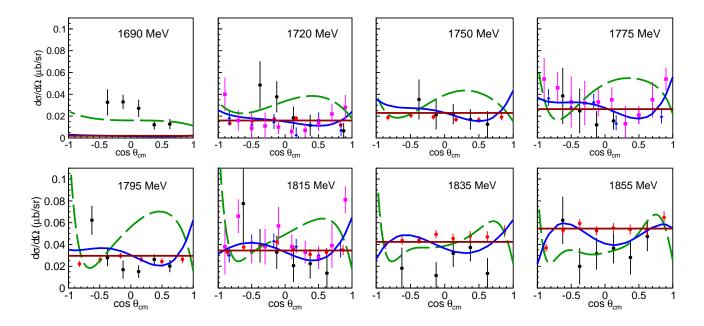


FIG. 6. Differential cross section for $\gamma p \to K^0 \Sigma^+$ for the various c.m. energy bins. The solid black circles show our results, the solid magenta squares show prior results from Lawall *et al.* [7], the solid blue triangles show prior results from Castelijns *et al.* [8], the solid red circles show prior results from Aguar-Bartolomé *et al.* [9], and the dashed green curves represent isobar-model predictions by Mart [29]. The solid blue curves show results of a 15-parameter global fit to our results and prior differential cross-section data. The solid red curves show results of a three-parameter global fit in which the angular distributions were approximated as being isotropic in each energy bin. (See text for details.)

The $a_{n\ell}$ coefficients were constant fitting parameters. Uncertainties in the fitted cross sections were conservatively calculated as twice the difference between results of the 15-parameter global fit and a separate three-parameter global fit in which the angular distributions were approximated as being isotropic in each energy bin (only the a_{n0} coefficients were varied).

Our measured integrated cross sections for $\gamma p \to K^0 \Sigma^+$ were obtained by making one-parameter Legendre fits of our measured differential cross sections. They are shown in Fig. 7 as solid black circles. Prior results from Lawall et al. [7], Castelijns et al. [8], and Aguar-Bartolomé et al. [9] are shown as solid magenta squares, solid blue triangles, and solid red circles, respectively. The results of our 15-parameter global fit are shown as solid cyan circles. The experimental results are compared with Mart's isobar-model predictions [29] shown as a dashed green curve.

B.
$$\gamma n \to K^0 \Lambda$$

Since the measured $\gamma p \to K^0 \Sigma^+$ cross sections for case (61) were imprecise due to low statistics and the low acceptance at backward and forward angles, the fitted world values of $(d\sigma/d\Omega)_{\gamma p \to K^0 \Sigma^+}$ and the measured K^0 yields for case (60) were used to calculate $\gamma n \to K^0 \Lambda$ cross sections for case (60). Because the

associated uncertainties in the fitted world values were relatively large at $\cos\theta_{\rm cm}=\pm0.875,$ those angle bins were excluded for all three K^0 photoproduction reactions. The c.m. energy range W = 1615 to 1765 MeV was divided into five bins of width 30 MeV, and the c.m. energy range W = 1765 to 1865 MeV was divided into five bins of width 20 MeV. The first two c.m. energy bins W = 1630 and 1660 MeV were below $\gamma p \to K^0 \Sigma^+$ threshold 1687 MeV. Therefore only $\gamma n \to K^0 \Lambda$ events can contribute to these bins. Figure 8 shows the differential cross section for $\gamma n \to K^0 \Lambda$ for these ten energy bins. Solid black circles show our results (weighted average of cases (60) and (70)). The solid magenta triangles and solid blue triangles respectively show the g10 and g13 results from Compton et al. [10] measured at JLab. Our results agree, within uncertainties, with the JLab g10 results in the energy bins at 1720 and 1835 MeV and with the JLab g13 results in the energy bin at 1855 MeV. Our results are similar in shape to the JLab g13 measurements at 1660, 1690, 1750, and 1795 MeV but are smaller in magnitude. It should be noted that the g10 and g13 results, where they overlap, are consistent for c.m. energies above about 1800 MeV, but the g13 results below that energy are all larger (especially at forward angles) than the g10 result that falls into our energy bin at 1690 MeV. The solid red curves in Fig. 8 show results of two-parameter Legendre polynomial fits to our measurements. The solid green curves show predictions based

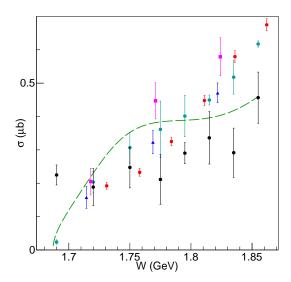


FIG. 7. Integrated cross section for $\gamma p \to K^0 \Sigma^+$. Our results, shown as solid black circles, were obtained by fitting our measured differential cross sections with a one-parameter Legendre expansion. The solid magenta squares show prior results from Lawall et al. [7], the solid blue triangles show prior results from Castelijns et al. [8], the solid red circles show prior results from Aguar-Bartolomé et al. [9], and the dashed green curve represents an isobar-model prediction by Mart [29]. The solid cyan circles were obtained from a 15-parameter global fit of our results combined with prior differential cross-section data. (See text for details.)

upon a partial-wave analysis [30]. Our results are in fair agreement with the predictions in all energy bins except at $1690~{\rm MeV}$.

We have checked various factors that might affect the normalizations of our results (e.g., the photon flux N_{γ} and detector acceptance) and have been unable to find any problems that would explain the differences between our results and the low-energy g13 results. Our results for all energy bins were handled in exactly the same manner as each other. Figure 9 shows the differential cross section for $\gamma n \to K^0 \Lambda$ as a function of c.m. energy W for individual angle bins. The results in this plot show a generally smooth energy variation, which implies we do not have normalization inconsistencies in individual energy bins.

Measured integrated cross sections for $\gamma n \to K^0 \Lambda$ are shown in Fig. 10. Solid black circles show our results, which were obtained by making two-parameter Legendre fits of the weighted average of our measured $\gamma n \to K^0 \Lambda$ differential cross sections for cases (60) and (70). The solid magenta triangles and solid blue triangles respectively show the g10 and g13 results from Compton *et al.* [10] measured at JLab. The solid green curve shows a prediction based upon a partial-wave analysis [30].

C.
$$\gamma n \to K^0 \Sigma^0$$

Our measured $\gamma n \to K^0 \Lambda$ differential cross sections for case (60) and our measured K^0 yields for case (70) were used to calculate the $\gamma n \to K^0 \Sigma^0$ differential cross sections for case (70). The c.m. energy range W = 1675 to 1765 MeV was divided into three bins of width 30 MeV, and the c.m. energy range W = 1765 to 1865 MeV was divided into five bins of width 20 MeV. Figure 11 shows the differential cross section for $\gamma n \to K^0 \Sigma^0$ (weighted average of cases (70) and (80)) for these eight c.m. energy bins. Our results are compared with isobar-model predictions (dashed blue curves) by Mart [29] and the solid red curves show results of two-parameter Legendre polynomial fits to our measurements. Our differential cross section results are in good agreement within error bars with Mart's predictions except at the highest energy bin, W = 1855 MeV. Figure 12 shows the differential cross section for $\gamma n \to K^0 \Sigma^0$ as a function of c.m. energy W for individual angle bins. As for $\gamma n \to K^0 \Lambda$, these results show a generally smooth energy variation, which supports the fact that the normalizations were determined consistently for the different energy bins.

Our measured integrated cross section values for $\gamma n \to K^0 \Sigma^0$ are shown in Fig. 13 as solid black circles. Our integrated cross sections were obtained by calculating the weighted average of our differential cross sections for cases (70) and (80) and then making two-parameter Legendre fits. Our experimental results are compared with an isobar-model prediction (solid blue curve) by Mart [29]. Our results are in good agreement with Mart's predictions except at the highest energy. These are the first experimental results for $\gamma n \to K^0 \Sigma^0$. As in the case of the differential cross sections, our results are in good agreement with Mart's predictions except at the highest energy bin.

VI. SUMMARY AND CONCLUSIONS

Our results for $\gamma p \to K^0 \Sigma^+$ are in fair agreement with prior measurements in the $\cos \theta_{\rm cm}$ range from -0.45 to +0.6, but our results at the most forward and backward angles are unreliable. For this reason, we used $\gamma p \to K^0 \Sigma^+$ world data to extract the $\gamma n \to K^0 \Lambda$ cross section for case (60). An isobar-model prediction by Mart [29] generally disagrees with all the measured differential cross sections.

Only one published set of prior measurements for $\gamma n \to K^0 \Lambda$ was available for comparing with our results. These prior results were measured with the CLAS spectrometer at JLab [10] in two separate datasets. In the seven energy bins where our results can be compared, our results agree within uncertainties with the g10 results but our results have a somewhat similar shape, but smaller magnitude, compared with the g13 results below W=1800 MeV. The results presented in Ref. [10] show that the g10 and g13 results, where they overlap,

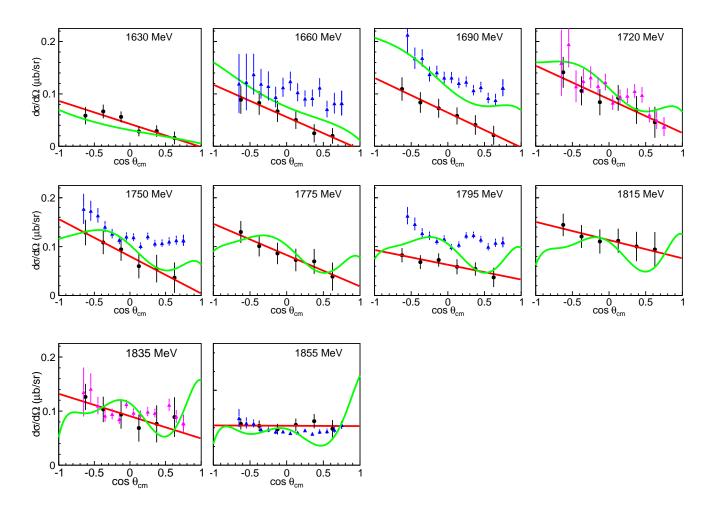


FIG. 8. Differential cross section for $\gamma n \to K^0 \Lambda$. The solid black circles represent the weighted average of our results for cases (60) and (70). The solid magenta triangles and solid blue triangles respectively show g10 and g13 results from Compton *et al.* [10]. The solid green curves are a prediction [30] based upon a partial-wave analysis. The solid red curves show the results of two-parameter Legendre polynomial fits to our measurements.

are generally consistent above about W=1800 MeV but not at lower energies. Our results are in fairly good agreement (except at 1690 MeV) with a prediction based on a partial-wave analysis [30] within error bars. Our results for $\gamma n \to K^0 \Lambda$ provide new measurements in the c.m. energy range from threshold (1614 MeV) to 1855 MeV.

Our results for $\gamma n \to K^0 \Sigma^0$ are the first experimental results for that reaction and span the c.m. energy range from the threshold (1691 MeV) to 1855 MeV. Our differential cross sections for $\gamma n \to K^0 \Sigma^0$ are in good agreement within error bars with isobar-model predictions by Mart [29] except at the highest energy bin. Our two independent measurements for cases (70) and (80) are consistent within error bars.

In summary, our new cross-section measurements for $\gamma n \to K^0 \Lambda$ and $\gamma n \to K^0 \Sigma^0$ will provide valuable data for future partial-wave analyses and will help better determine the properties of N^* resonances that decay to

 $K\Lambda$ or $K\Sigma$ final states.

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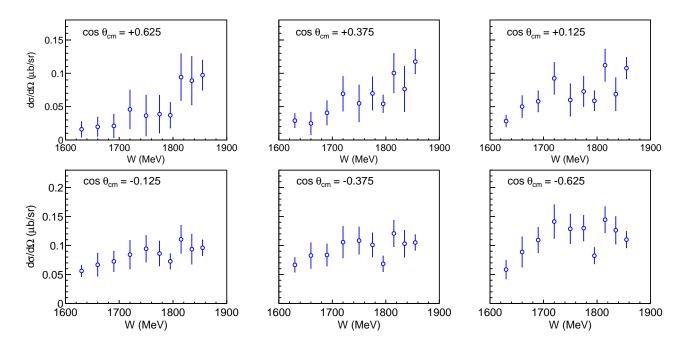


FIG. 9. Differential cross section for $\gamma n \to K^0 \Lambda$ versus c.m. energy W for angle bins from $\cos \theta_{\rm cm} = +0.625$ to -0.625. The open blue circles represent the weighted average of our results for cases (60) and (70).

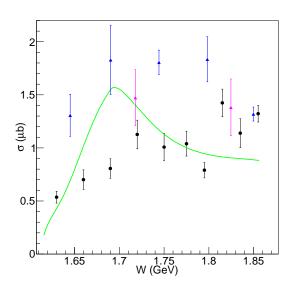


FIG. 10. Integrated cross section for $\gamma n \to K^0 \Lambda$. The solid black circles represent our results. The solid magenta triangles and solid blue triangles respectively show the g10 and g13 results from Compton *et al.* [10]. The solid green curve shows a prediction [30] based upon a partial-wave analysis.

nology Facilities Council(STFC 57071/1, 50727/1), European Community Research Infrastructure Activity (FP6), the U.S. DOE, U.S. NSF, and NSERC (Canada).

W. J. Briscoe et al., Eur. Phys. J A 51, 129 (2015); and references therein.

^[2] R. G. T. Zegers et~al., Phys. Rev. Lett. ${\bf 91},\,092001$ (2003).

^[3] J. W. C. McNabb et al. (CLAS Collaboration), Phys.

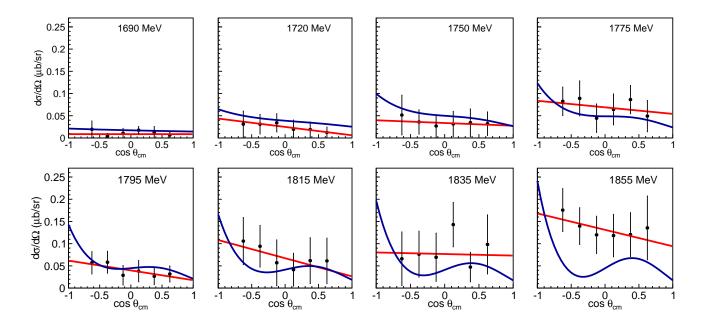


FIG. 11. Differential cross section for $\gamma n \to K^0 \Sigma^0$. Solid black circles show our results. Solid blue curves represent isobar-model predictions by Mart [29] and the solid red curves show results of two-parameter Legendre polynomial fits to our measurements.

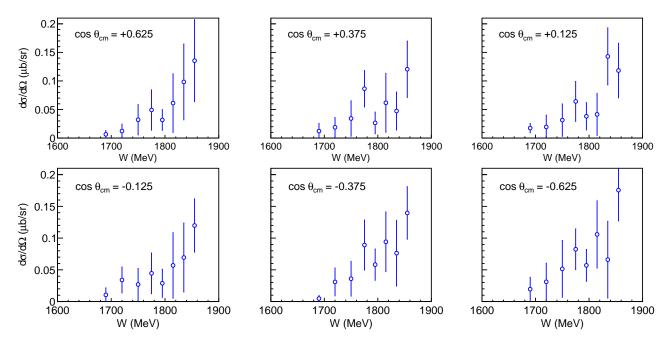


FIG. 12. Differential cross section for $\gamma n \to K^0 \Sigma^0$ versus c.m. energy W for angle bins from $\cos \theta_{\rm cm} = +0.625$ to -0.625. The open blue circles represent the weighted average of our results for cases (70) and (80).

Rev. C **69**, 042201 (2004).

- [4] M. Q. Tran *et al.* (SAPHIR Collaboration), Phys. Lett. B **445**, 20 (1998).
- [5] R. Bradford *et al.* (CLAS Collaboration), Phys. Rev. C 73, 035202 (2006).
- [6] A. V. Sarantsev et al., Eur. Phys. J. A 25, 441 (2005).
- [7] R. Lawall et al., Eur. Phys. J. A 24, 275 (2005).
- [8] R. Castelijns et al. (The CBELSA/TAPS Collaboration), Eur. Phys. J. A 35, 39 (2008).
- [9] P. Aguar-Bartolomé et al. (A2 Collaboration at MAMI), Phys. Rev. C 88, 044601 (2013).
- [10] N. Compton et al. (CLAS Collaboration), Phys. Rev. C

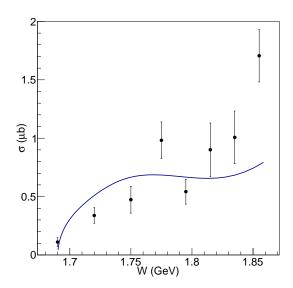


FIG. 13. Integrated cross section for $\gamma n \to K^0 \Sigma^0$. Solid black circles show our results. The solid blue curve represents an isobar-model prediction by Mart [29].

96, 065201 (2017).

- [11] A. Starostin et al. (The Crystal Ball Collaboration), Phys. Rev. C 64, 055205 (2001).
- [12] D. Werthmüller, Ph.D. Thesis, University of Basel (2014).
- [13] S. Prakhov *et al.* (Crystal Ball Collaboration at MAMI and A2 Collaboration), Phys. Rev. C **79**, 035204 (2009).
- [14] E. F. McNicoll et al. (Crystal Ball Collaboration at MAMI), Phys. Rev. C 82, 035208 (2010).
- [15] M. Dieterle, et al. (A2 Collaboration), Phys. Rev. C $\bf 97$, 065205 (2018).

- [16] D. Watts, Calorimetry in Particle Physics, Proceedings of the 11th International Conference, Perugia, Italy, 2004 (World Scientific, 2005), p. 560.
- [17] K.-H. Kaiser et al., Nucl. Instr. Meth. A 593, 159 (2008).
- [18] J. C. McGeorge et al., Eur. Phys. J. A 37, 129 (2008).
- [19] K. R. Bantawa, Ph.D. Thesis, Kent State University (2009).
- [20] D. Werthmüller, et al. (A2 Collaboration at MAMI), Phys. Rev. C 90, 015205 (2014).
- [21] L. Witthauer et al. (The A2 Collaboration), Eur. Phys. J. A 49, 154 (2013).
- [22] F. Zehr et al. (The Crystal Ball at MAMI, TAPS, and A2 Collaborations), Eur. Phys. J. A 48, 98 (2012).
- [23] A. Käser et al. (The A2 Collaboration), Eur. Phys. J. A 52, 272 (2016).
- [24] M. Oberle et al., Phys. Lett. B 721, 237 (2013).
- [25] S. Schumann et al. (The Crystal Ball at MAMI, TAPS, and A2 Collaborations), Eur. Phys. J. A 43, 269 (2010).
- [26] M. Nanova et al. (The CBELSA/TAPS Collaboration), Eur. Phys. J. A 35, 333 (2008).
- [27] M. Tanabashi *et al.* (Particle Data Group), Phys. Rev. D 98, 030001 (2018).
- [28] ChandraSekhar Akondi, Ph.D. Thesis, Kent State University (2018).
- [29] T. Mart Phys. Rev. C 90, 065202 (2014) and private communication (2017).
- [30] B. C. Hunt, Ph.D. Thesis, Kent State University (2017);
 B. C. Hunt and D. M. Manley, arXiv:1804.07422 [nuclex], (2018).

Appendix: Tabulation of Results

In this appendix, we provide our measured differential and integrated cross sections for $\gamma p \to K^0 \Sigma^+$ in Tables II and III, our measured differential and integrated cross sections for $\gamma n \to K^0 \Lambda$ in Tables IV and V, and our measured differential and integrated cross sections for $\gamma n \to K^0 \Sigma^0$ in Tables VI and VII.

TABLE II. Differential cross section for $\gamma p \to K^0 \Sigma^+$. Systematic uncertainties less than 0.001 are not listed.

\overline{W}	$\cos \theta_{ m cm}$	$d\sigma/d\Omega$	stat. unc.	sys. unc.	\overline{W}	$\cos \theta_{ m cm}$	$d\sigma/d\Omega$	stat. unc.	sys. unc.
(MeV)	COBVEIN	$(\mu b/sr)$	$(\mu \mathrm{b/sr})$	$(\mu b/sr)$	(MeV)	COBUCIII	$(\mu b/sr)$	$(\mu b/sr)$	$(\mu \mathrm{b/sr})$
1690	+0.625	0.013	0.004	0.001	1795	+0.625	0.020	0.005	0.001
1690	+0.375	0.012	0.004	0.001	1795	+0.375	0.026	0.004	0.001
1690	+0.125	0.027	0.007	0.002	1795	+0.125	0.015	0.005	0.001
1690	-0.125	0.033	0.006	0.002	1795	-0.125	0.017	0.008	0.001
1690	-0.375	0.033	0.012	0.002	1795	-0.375	0.028	0.007	0.001
1720	+0.625	_	_	_	1815	+0.625	0.014	0.014	0.001
1720	+0.375	0.012	0.009	0.001	1815	+0.375	0.022	0.013	0.002
1720	+0.125	0.018	0.010	0.001	1815	+0.125	0.021	0.015	0.001
1720	-0.125	0.038	0.014	0.002	1815	-0.125	0.033	0.015	0.002
1720	-0.375	0.049	0.021	0.002	1815	-0.375	0.039	0.015	0.003
1750	+0.625	0.012	0.012	_	1835	+0.625	0.013	0.014	0.001
1750	+0.375	0.017	0.009	0.001	1835	+0.375	0.037	0.012	0.003
1750	+0.125	0.022	0.009	0.001	1835	+0.125	0.032	0.012	0.003
1750	-0.125	0.021	0.013	0.001	1835	-0.125	0.012	0.012	0.001
1750	-0.375	0.035	0.018	0.001	1835	-0.375	0.003	0.003	_
1775	+0.625	0.005	0.005	_	1855	+0.625	0.047	0.013	0.004
1775	+0.375	0.006	0.006	_	1855	+0.375	0.028	0.015	0.002
1775	+0.125	0.016	0.008	0.001	1855	+0.125	0.037	0.013	0.003
1775	-0.125	0.012	0.012	0.001	1855	-0.125	0.032	0.014	0.003
1775	-0.375	0.025	0.020	0.001	1855	-0.375	0.020	0.016	0.002

TABLE III. Integrated cross section for $\gamma p \to K^0 \Sigma^+.$

\overline{W}	σ	stat. unc.	sys. unc.
(MeV)	(μb)	$(\mu \mathrm{b})$	$(\mu \mathrm{b})$
1690	0.225	0.030	0.010
1720	0.188	0.055	0.006
1750	0.247	0.061	0.005
1775	0.211	0.075	0.007
1795	0.290	0.031	0.011
1815	0.336	0.079	0.023
1835	0.291	0.074	0.023
1855	0.457	0.077	0.035

TABLE IV. Differential cross section for $\gamma n \to K^0 \Lambda.$

W	$\cos \theta_{ m cm}$	$d\sigma/d\Omega$	stat. unc.	sys. unc.	W	$\cos heta_{ m cm}$	$d\sigma/d\Omega$	stat. unc.	sys. unc.
(MeV)		$(\mu \mathrm{b/sr})$	$(\mu \mathrm{b/sr})$	$(\mu \mathrm{b/sr})$	(MeV)		$(\mu \mathrm{b/sr})$	$(\mu \mathrm{b/sr})$	$(\mu \mathrm{b/sr})$
1630	+0.625	0.016	0.011	0.001	1775	+0.625	0.039	0.028	0.002
1630	+0.375	0.029	0.010	0.001	1775	+0.375	0.070	0.025	0.003
1630	+0.125	0.028	0.009	0.001	1775	+0.125	0.073	0.022	0.003
1630	-0.125	0.056	0.010	0.002	1775	-0.125	0.086	0.021	0.004
1630	-0.375	0.066	0.012	0.003	1775	-0.375	0.101	0.020	0.004
1630	-0.625	0.058	0.016	0.002	1775	-0.625	0.130	0.022	0.005
1660	+0.625	0.020	0.014	0.001	1795	+0.625	0.037	0.019	0.002
1660	+0.375	0.025	0.017	0.001	1795	+0.375	0.054	0.013	0.002
1660	+0.125	0.050	0.016	0.002	1795	+0.125	0.059	0.015	0.003
1660	-0.125	0.067	0.020	0.003	1795	-0.125	0.073	0.013	0.003
1660	-0.375	0.083	0.022	0.004	1795	-0.375	0.068	0.013	0.003
1660	-0.625	0.089	0.026	0.004	1795	-0.625	0.082	0.014	0.004
1690	+0.625	0.021	0.017	0.001	1815	+0.625	0.094	0.035	0.007
1690	+0.375	0.041	0.018	0.002	1815	+0.375	0.100	0.029	0.007
1690	+0.125	0.058	0.016	0.003	1815	+0.125	0.112	0.024	0.008
1690	-0.125	0.073	0.017	0.004	1815	-0.125	0.111	0.024	0.008
1690	-0.375	0.084	0.019	0.004	1815	-0.375	0.121	0.023	0.009
1690	-0.625	0.109	0.022	0.005	1815	-0.625	0.144	0.022	0.010
1720	+0.625	0.046	0.029	0.002	1835	+0.625	0.089	0.036	0.007
1720	+0.375	0.069	0.026	0.003	1835	+0.375	0.076	0.034	0.006
1720	+0.125	0.092	0.024	0.004	1835	+0.125	0.069	0.024	0.006
1720	-0.125	0.084	0.024	0.003	1835	-0.125	0.094	0.026	0.008
1720	-0.375	0.106	0.027	0.004	1835	-0.375	0.103	0.023	0.008
1720	-0.625	0.141	0.029	0.005	1835	-0.625	0.126	0.023	0.010
1750	+0.625	0.036	0.030	0.001	1855	+0.625	0.097	0.022	0.008
1750	+0.375	0.055	0.027	0.002	1855	+0.375	0.118	0.018	0.010
1750	+0.125	0.060	0.024	0.002	1855	+0.125	0.108	0.016	0.009
1750	-0.125	0.094	0.023	0.003	1855	-0.125	0.096	0.013	0.008
1750	-0.375	0.108	0.023	0.003	1855	-0.375	0.105	0.013	0.008
1750	-0.625	0.129	0.025	0.003	1855	-0.625	0.110	0.014	0.009

TABLE V. Integrated cross section for $\gamma n \to K^0 \Lambda$.

\overline{W}	σ	stat. unc.	sys. unc.
(MeV)	(μb)	$(\mu \mathrm{b})$	$(\mu \mathrm{b})$
1630	0.54	0.05	0.02
1660	0.70	0.09	0.03
1690	0.81	0.09	0.03
1720	1.13	0.13	0.02
1750	1.01	0.13	0.04
1775	1.04	0.12	0.04
1795	0.79	0.07	0.05
1815	1.42	0.13	0.11
1835	1.14	0.13	0.09
1855	1.32	0.08	0.02

TABLE VI. Differential cross section for $\gamma n \to K^0 \Sigma^0$. Systematic uncertainties less than 0.001 are not listed.

W	$\cos \theta_{ m cm}$	$d\sigma/d\Omega$	stat. unc.	sys. unc.	W	$\cos \theta_{ m cm}$	$d\sigma/d\Omega$	stat. unc.	sys. unc.
(MeV)		$(\mu \mathrm{b/sr})$	$(\mu \mathrm{b/sr})$	$(\mu \mathrm{b/sr})$	(MeV)		$(\mu \mathrm{b/sr})$	$(\mu \mathrm{b/sr})$	$(\mu \mathrm{b/sr})$
1690	+0.625	0.007	0.007	_	1795	+0.625	0.032	0.018	0.002
1690	+0.375	0.012	0.013	0.001	1795	+0.375	0.027	0.019	0.001
1690	+0.125	0.017	0.008	0.001	1795	+0.125	0.038	0.024	0.002
1690	-0.125	0.011	0.011	0.001	1795	-0.125	0.029	0.022	0.001
1690	-0.375	0.005	0.005	_	1795	-0.375	0.058	0.025	0.003
1690	-0.625	0.020	0.019	0.001	1795	-0.625	0.057	0.025	0.003
1720	+0.625	0.012	0.012	0.001	1815	+0.625	0.061	0.051	0.005
1720	+0.375	0.019	0.017	0.001	1815	+0.375	0.062	0.052	0.005
1720	+0.125	0.020	0.021	0.001	1815	+0.125	0.042	0.037	0.003
1720	-0.125	0.034	0.020	0.002	1815	-0.125	0.057	0.052	0.004
1720	-0.375	0.031	0.022	0.002	1815	-0.375	0.094	0.047	0.007
1720	-0.625	0.031	0.029	0.002	1815	-0.625	0.106	0.053	0.008
1750	+0.625	0.032	0.027	0.001	1835	+0.625	0.098	0.066	0.008
1750	+0.375	0.034	0.031	0.002	1835	+0.375	0.047	0.033	0.004
1750	+0.125	0.031	0.028	0.001	1835	+0.125	0.143	0.050	0.012
1750	-0.125	0.027	0.025	0.001	1835	-0.125	0.069	0.054	0.006
1750	-0.375	0.036	0.027	0.002	1835	-0.375	0.076	0.052	0.007
1750	-0.625	0.052	0.045	0.002	1835	-0.625	0.066	0.060	0.006
1775	+0.625	0.049	0.035	0.002	1855	+0.625	0.136	0.072	0.012
1775	+0.375	0.086	0.032	0.005	1855	+0.375	0.120	0.049	0.011
1775	+0.125	0.064	0.035	0.003	1855	+0.125	0.118	0.048	0.010
1775	-0.125	0.045	0.032	0.002	1855	-0.125	0.120	0.042	0.010
1775	-0.375	0.089	0.039	0.005	1855	-0.375	0.140	0.041	0.012
1775	-0.625	0.082	0.032	0.004	1855	-0.625	0.176	0.048	0.015

TABLE VII. Integrated cross section for $\gamma n \to K^0 \Sigma^0$.

\overline{W}	σ	stat. unc.	sys. unc.
(MeV)	(μb)	$(\mu \mathrm{b})$	(μb)
1690	0.111	0.038	0.005
1720	0.338	0.067	0.011
1750	0.47	0.12	0.01
1775	0.98	0.16	0.03
1795	0.54	0.11	0.02
1815	0.90	0.23	0.06
1835	1.01	0.23	0.08
1855	1.71	0.22	0.13